Hartree-Fock-Bogoliubov (HFB) theory

General remarks

HF + BCS theory: mean field and pairing are decoupled.
 BCS pairing is added to HF mean field calculations (fairly trivial modification: replace HF occupation numbers by those of BCS).
 This approach works well near the valley of stability.

• HFB theory: mean field and pairing field influence each other. Self-consistent theory of both phenomena. Start from generalized variational principle. Canonical transformation (Bogoliubov) to "quasi-particles".

This approach is essential for nuclei far away from stability.

HFB formalism: quasi-particle transformation

Creation operators for nucleon states: \hat{c}_i^{\dagger}

$$\hat{H} = \sum_{i,j=1}^{\infty} \langle i|t|j \rangle \hat{c}_{i}^{\dagger} \hat{c}_{j} + \frac{1}{2} \sum_{i,j,k,l=1}^{\infty} \langle ij|V^{(2)}|kl \rangle \hat{c}_{i}^{\dagger} \hat{c}_{j}^{\dagger} \hat{c}_{l} \hat{c}_{k}$$
$$+ \frac{1}{6} \sum_{i,j,k,l,m,n=1}^{\infty} \langle ijk|V^{(3)}|lmn \rangle \hat{c}_{i}^{\dagger} \hat{c}_{j}^{\dagger} \hat{c}_{k}^{\dagger} \hat{c}_{n} \hat{c}_{m} \hat{c}_{l}$$

Canonical transformation to quasi-particles: $\hat{\beta}_{k}$

ground state = quasiparticle vacuum

$$\hat{\beta}_{k}|\Phi_{0}
angle=0$$

HFB energy density functional and generalized variational principle



HFB equations for quasi-particle wave functions (in coordinate space)

$$\begin{pmatrix} (h-\lambda) & \tilde{h} \\ \tilde{h} & -(h-\lambda) \end{pmatrix} \begin{pmatrix} \phi_{\alpha}^{(1)} \\ \phi_{\alpha}^{(2)} \\ \phi_{\alpha}^{(2)} \end{pmatrix} = E_{\alpha} \begin{pmatrix} \phi_{\alpha}^{(1)} \\ \phi_{\alpha}^{(2)} \\ \phi_{\alpha}^{(2)} \end{pmatrix}$$

Mean-field Hamiltonian

Pairing-field Hamiltonian

$$h = \begin{pmatrix} h_{\uparrow\uparrow}(\vec{r}) & h_{\uparrow\downarrow}(\vec{r}) \\ h_{\downarrow\uparrow}(\vec{r}) & h_{\downarrow\downarrow}(\vec{r}) \end{pmatrix}$$

$$\tilde{h} = \begin{pmatrix} \tilde{h}_{\uparrow\uparrow}(\vec{r}) & \tilde{h}_{\uparrow\downarrow}(\vec{r}) \\ \tilde{h}_{\downarrow\uparrow}(\vec{r}) & \tilde{h}_{\downarrow\downarrow}(\vec{r}) \end{pmatrix}$$

 $\phi_{\alpha}^{(2)} = \begin{pmatrix} \phi_{\alpha}^{(2)}(\vec{r},\uparrow) \\ \phi_{\alpha}^{(2)}(\vec{r},\downarrow) \end{pmatrix}$

two types of spinor wave functions

$$\boldsymbol{\phi}_{\alpha}^{(1)} = \begin{pmatrix} \boldsymbol{\phi}_{\alpha}^{(1)}(\vec{r},\uparrow) \\ \boldsymbol{\phi}_{\alpha}^{(1)}(\vec{r},\downarrow) \end{pmatrix}$$

HFB particle density and pairing density in terms of quasi-particle wave functions

$$\begin{split} \rho_{q}(\vec{r}) &= \sum_{\alpha} \sum_{\sigma=\uparrow\downarrow} \left| \phi_{\alpha}^{(2)}(\vec{r} \, q \, \sigma) \right|^{2} & \text{normal density} \\ \tilde{\rho}_{q}(\vec{r}) &= -\sum_{\alpha} \sum_{\sigma=\uparrow\downarrow} \phi_{\alpha}^{(2)}(\vec{r} \, q \, \sigma) \phi_{\alpha}^{(1)*}(\vec{r} \, q \, \sigma) & \text{pairing density} \end{split}$$

Observables predicted by HFB theory

Same as in HF theory, but in addition:

- Pairing density distribution of protons and neutrons
- \Rightarrow pairing energy for p / n
- \Rightarrow average pairing gap for p / n
- quasi-particle energy spectrum (and equivalent s.p. energy spectrum)

NOTE: HFB is a ground state theory. For excited states we need Quasi-particle Random Phase Approximation = QRPA

HFB codes in coordinate space

- 1-D radial grid (for spherical nuclei)
 Dobaczewski, Flocard & Treiner, Nucl. Phys. A422, 103 (1984)
- 3-D Cartesian lattice, diagonalize HFB in HF basis maximum continuum energy is quite small (5 MeV) Terasaki, Heenen, Flocard & Bonche, Nucl. Phys. A600, 371 (1996)
- 2-D grid (axial), HFB in "transformed harmonic oscillator" (THO) basis Stoitsov, Dobaczewski, Ring & Pittel, Phys. Rev. C61, 034311 (2000)
- 2-D grid (axial), HFB in B-Spline representation
 VU group: Oberacker, Umar, Teran, Blazkiewicz (2003-2007)
- 2-D grid (axial), HFB in B-Spline representation Pei, Stoitsov, Fann, Nazarewicz, Schunck & Xu, Phys. Rev. C78, 064306 (2008)

Comparison of calculated binding energies with experimental data

HF / HFB calculations use Skyrme (SLy4) interaction + delta pairing interaction

Binding energies [MeV]			
Nucleus	HF + BCS/LN (2D)	HFB (2D)	Exp.
²⁴ O	-173.76	-172.47	-168.96
⁹⁶ Zr	-828.44	-826.91	-829.00
¹²⁴ Sn	-1050.39	-1049.75	-1049.96

Note: none of the above nuclei are used in the Skyrme (SLy4) force fit!

Stable nucleus ¹⁶O HFB-2D calculation, SLy4 N-N interaction



Neutron dripline nucleus $^{24}O(T_{1/2} = 65 \text{ ms})$ HFB-2D calculation, SLy4 N-N interaction



2n-dripline nucleus ¹²²Zr (Z=40, N=82), HFB-2D, SLy4



Neutrons at Fermi energy are almost unbound. Neutron wave functions spread out much more than those of protons: "neutron skin"

HFB in coordinate space (2-D / 3-D lattice)

Computational Challenges near neutron-dripline

3 types of wavefunctions

- well-bound states
- weakly-bound states
- continuum states

Other features

- strong pairing correlations for nuclei near neutron dripline
- strong continuum coupling

For nuclei near the neutron dripline, we need accurate solution of HFB continuum problem on coordinate-space lattice, for single-particle energies up to 60 MeV, or even higher for collective excited states via QRPA

Strong continuum coupling near the neutron dripline Matsuo & Nakatsukasa, Journal of Physics G, May 2010



Figure 2. Schematic picture depicting correlations in nuclei near the drip-line, for which correlations involving the loosely bound and unbound continuum orbits play essential roles.

New physics near the neutron dripline

- strong pairing correlations (HF+BCS becomes unrealistic, need full HFB)
- highly asymmetric Fermi levels (-20 MeV for protons, -1 MeV for neutrons)
- weakly bound neutron states near Fermi level neutron halos / neutron skins → neutron star physics! strong continuum coupling of weakly bound neutrons
- reduced LS coupling → disappearance of shell gaps
 → abundance of elements created via r process (Supernovae)
- new collective modes ("pygmy" resonance, "scissors" vibrations, ...)

¹⁰²Zr: normal density and pairing density HFB, 2-D lattice, SLy4 + volume pairing Artur Blazkiewicz, Vanderbilt, Ph.D. thesis (2005)



HFB: $\beta_2^{(p)}=0.43$

exp: $\beta_2^{(p)}=0.42(5)$, J.K. Hwang et al., Phys. Rev. C (2006)

¹¹⁰Zr (Z=40, N=70), HFB on 2-D grid Blazkiewicz, Oberacker, Umar & Stoitsov, Phys. Rev. C71, 054321 (2005)



¹⁰⁴Zr (Z=40, N=64), HFB on 2D lattice neutron pairing density: single-particle energy spectrum Oberacker et al., Proc. NENS03 conference (Niigata, Japan, Nov. 2003)

Zirconium isotopes (Z=40), HFB+SLy4 on 2D-grid Blazkiewicz, Oberacker, Umar & Stoitsov, Phys. Rev. C71, 054321 (2005)

2n-dripline nucleus ¹²²Zr (Z=40, N=82), HFB on 2-D grid Blazkiewicz, Oberacker, Umar & Stoitsov, Phys. Rev. C71, 054321 (2005)

Average neutron pairing gap for zirconium isotopes (Z=40) HFB-2D (SLy4 + volume pairing)

Zirconium isotopes (Z=40), HFB+SLy4 on 2D-grid Blazkiewicz, Oberacker, Umar & Stoitsov, Phys. Rev. C71, 054321 (2005)

Tin isotopes (Z=50), HFB-1D for several Skyrme interactions

RIA Theory Bluebook (Sep. 2005), calculations by Dobaczewski et al.

2-neutron separation energies (2-D HFB)

Dobaczewski, Stoitsov & Nazarewicz (2004) arXiv:nucl-th/0404077

Nuclear ground state deformations (2-D HFB)

Dobaczewski, Stoitsov & Nazarewicz (2004) arXiv:nucl-th/0404077

HFB with quadrupole + octupole constraints H. Goutte, P. Casoli, and J.F. Berger, Nucl. Phys. A734 (2004) 217

microscopic potential energy surface $E(Q_{20}, Q_{30})$ for nuclear fission: Q_{20} – elongation Q_{30} – mass asymmetry

Notice double-humped fission barrier in Q_{20} – direction !

HFB with quadrupole + octupole constraints Robledo, Baldo, Schuck, and Viñas, Phys. Rev. C 81, 034315 (2010)

single-particle energies for proton levels in ²²⁴Ra, as a function of Q_2 – mass quadrupole moment Q_3 – octupole moment Total g.s. energy minimum at Q_2 = 8.12 b, Q_3 = 4 b^{3/2}

Constrained HFB calculation with Gogny effective N-N interaction.

- Fermi level

Quenching of nuclear shell structure in neutron-rich nuclei

near stability line: pronounced shell gaps n-rich nuclei : shell gaps suppressed due to reduced LS coupling

Solar r-process: cosmic element abundance RIA 2000 Whitepaper

